## Problem Solving Physics Toolkit

Most of this should be familiar to you from A level physics, but as the different A levels vary somewhat you may find one or two things here that you don't know. All the examples we use can be solved with just these equations (plus any others you may deduce from the problem itself, of course).

## Newton's Laws and Mechanics

| Newton's second law | $F=m a$ | $F=$ force, $m=$ mass, <br> $a=$ acceleration |
| :--- | :---: | :--- |
| Impulse (integrated force <br> applied over short time) | $F \Delta t=\Delta p$ | $F=$ force, $\Delta t=$ time interval, <br> $\Delta p=$ change in momentum |
| Newton's law of gravity | $a=\frac{G m_{1} m_{2}}{r^{2}}$ | $F=$ force, $G=$ gravitational <br> constant, $m_{1}, m_{2}=$ masses, <br> $r=$ separation |
| Uniform circular motion | $F=\mu N$ | $a=$ acceleration, $v=$ speed of <br> motion of object moving in a <br> circle of radius $r, \omega=$ angular <br> velocity of object $=v / r$ |
| Friction | $F=k \Delta x$ | $F=$ force, $\mu=$ coefficient of <br> friction, $N=$ normal force; $F$ <br> is directed opposite to <br> direction of motion of object |
| Hooke's law | $F=$ force, $k=$ spring constant, <br> $\Delta x=$ change in spring length |  |

## ELECTROMAGNETISM

| Coulomb's law | $F=\frac{Q_{1} Q_{2}}{4 \pi \varepsilon_{0} r^{2}}$ | $F=$ force, $Q_{1}, Q_{2}=$ charges, $\varepsilon_{0}=$ <br> permittivity of free space <br> (constant), $r=$ separation |
| :--- | :--- | :--- |
| Magnetic force | $\mathbf{F}=q \mathbf{v} \times \mathbf{B}$ | $\mathbf{F}=$ force (vector), $q=$ charge, <br> $\mathbf{v}=$ velocity (vector), <br> $\mathbf{B}=$ magnetic field (vector) |
| Speed of light | $c^{2}=\frac{1}{\mu_{0} \varepsilon_{0}}$ | $c=$ speed of light, <br> $\mu_{0}=$ permeability of free space <br> (constant), $\varepsilon_{0}=$ permittivity of <br> free space (constant) |

The latter two equations will be used mostly in dimensional analysis problems.

## Thermal Physics

| Ideal gas law | $P V=N k_{\mathrm{B}} T$ | $P=$ pressure, $V=$ volume, <br> $N=$ number of particles, <br> $k_{\mathrm{B}}=$ Boltzmann's constant, <br> $T=$ absolute temperature |
| :--- | :--- | :--- |
| Blackbody radiation ${ }^{1}$ <br> (energy radiated by a body at <br> temperature $T$, assuming perfect <br> efficiency) | $f=\sigma T^{4}$ | $f=$ power radiated per unit <br> area (flux), $\sigma=$ Stefan's <br> constant, $T=$ absolute <br> temperature |

## Useful Mathematical Formulae

| Sum of $N$ terms of an arithmetic series whose $n^{\text {th }}$ term is $a_{n}=a_{1}+(n-1) k$ | $S_{N}=\sum_{i=0}^{N-1}\left(a_{1}+i k\right)=\frac{N}{2}\left(a_{1}+a_{N}\right)$ |
| :---: | :---: |
| Sum of $N$ terms of a geometric series whose $n^{\text {th }}$ term is $a_{n}=a_{1} f^{n-1}$ | $S_{N}=a_{1} \sum_{i=0}^{N-1} f^{i}=a_{1} \frac{1-f^{N}}{1-f}$ |
| Taylor series for $f(x)$ around $f(a)$ ( $f^{\prime}=\mathrm{d} f / \mathrm{d} x$ etc.) | $\begin{aligned} f(x)=f(a)+ & f^{\prime}(a)(x-a) \\ & +\frac{1}{2!} f^{\prime \prime}(a)(x-a)^{2} \\ & +\frac{1}{3!} f^{\prime \prime \prime}(a)(x-a)^{3}+\cdots \end{aligned}$ |
| Commonly used series expansions: | $\begin{aligned} & \sin \theta=\theta-\frac{1}{3!} \theta^{3}+\frac{1}{5!} \theta^{5}-\cdots(\theta \text { in radians }) \\ & \cos \theta=1-\frac{1}{2!} \theta^{2}+\frac{1}{4!} \theta^{4}-\cdots(\theta \text { in radians }) \\ & (1+x)^{n}=1+n x+\frac{1}{2} n(n-1) x^{2}+ \\ & \frac{1}{3!} n(n-1)(n-2) x^{3}+\cdots \\ & e^{x}=1+x+\frac{1}{2!} x^{2}+\frac{1}{3!} x^{3}+\cdots \\ & \ln (1+x)=x-\frac{1}{2} x^{2}+\frac{1}{3} x^{3}-\cdots(\text { for }\|x\|<1) \end{aligned}$ |

## Numerical Constants

| $G=6.67 \times 10^{-11} \mathrm{~N} \mathrm{~m}^{2} \mathrm{~kg}^{-2}$ | $k_{\mathrm{B}}=1.38 \times 10^{-23} \mathrm{~J} \mathrm{~K}^{-1}$ | $\varepsilon_{0}=8.85 \times 10^{-12} \mathrm{~F} \mathrm{~m}^{-1}$ |
| :--- | :--- | :--- |
| $g=9.81 \mathrm{~m} \mathrm{~s}^{-2}$ | $\sigma=5.67 \times 10^{-8} \mathrm{~W} \mathrm{~m}^{-2} \mathrm{~K}^{-4}$ | $\mu_{0}=4 \pi \times 10^{-7} \mathrm{H} \mathrm{m}^{-1}$ |
| Earth's mass $=6.0 \times 10^{24} \mathrm{~kg}$ | $N_{\mathrm{A}}=6.02 \times 10^{23}(\mathrm{~g}-\mathrm{mol})^{-1}$ | $c=2.998 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}$ |
| Sun's mass $=2.0 \times 10^{30} \mathrm{~kg}$ | 1 mole of gas: 22.4 litres $(\mathrm{STP})$ | $e=1.6 \times 10^{-19} \mathrm{C}$ |
| $h=6.63 \times 10^{-34} \mathrm{~J} \mathrm{~s}$ | $m_{\mathrm{e}}=9.1 \times 10^{-31} \mathrm{~kg}$ | $m_{\mathrm{p}}=1.67 \times 10^{-27} \mathrm{~kg}$ |

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[^0]:    ${ }^{1}$ This you probably didn't meet in A level, but it's useful, so we'll include it here.

