

PROBLEM SOLVING

Calculations

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- Using expansions
- Checks and estimates

CALCULATIONS

- Numerical value: last step in solving a problem
 - usually doesn't cost many exam marks if you get it wrong
 - provided method and logic correct
 - but obviously crucial in real-world applications
 - underestimating stresses on an aircraft wing by factor of 10 clearly not a good idea
- Usual rules apply
 - Know what you're doing
 - Simplify if possible
 - Check answer
 - "calculator typos" common cause of lost marks in exams

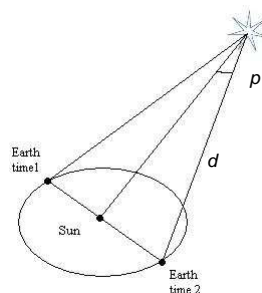
EXPANSIONS AND APPROXIMATIONS

- Solutions to many problems rely on series expansions
 - e.g. small angle approximations, $\sin \theta \approx \theta$, $\cos \theta \approx 1 - \frac{1}{2}\theta^2$
- Know the most useful ones
 - $\sin \theta$, $\cos \theta$, $(1 + x)^n \approx 1 + nx$ (binomial expansion)
 - they are on standard constants sheet if you forget
 - but often needed in derivations, so better to know them
- Be aware of any limitations
 - $\sin \theta$, $\cos \theta$ need θ in radians
- When to use in calculations
 - when required by derivation/definition
 - when they introduce simplifications
 - when you might otherwise have problems with precision

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EXAMPLE 1: DEFINITION

- The parallax of Sirius is 0.379 arc seconds. What is its distance in parsecs?



- The definition of the parsec is “the distance at which a star would have a parallax of one arc second” (and angles in arc seconds are always small)
 - so $d(\text{pc}) = 1/p(\text{arcsec})$ by definition, hence $d = 2.64 \text{ pc}$
- We could:
 - convert $0.379''$ to degrees: 0.0001053°
 - take the sine of that: 1.84×10^{-6}
 - look up the distance to the Sun in metres: $1.50 \times 10^{11} \text{ m}$
 - write $d = 1.50 \times 10^{11} / 1.84 \times 10^{-6} = 8.1 \times 10^{16} \text{ m}$
 - look up the conversion between parsecs and metres:
 $1 \text{ pc} = 3.09 \times 10^{16} \text{ m}$
 - and convert d to pc: $d = 8.1 / 3.09 = 2.64 \text{ pc}$
 - Same answer, six times more work!

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EXAMPLE 2: SIMPLIFICATION

- If you are on a boat in mid-ocean, how far away is the horizon? Assume that when standing on the deck your eye level is 10 m above the water surface, and take the Earth to be a sphere of radius 6370 km.
 - You want the short side of a right-angled triangle whose other two sides are 6370 km and (6370km + 10 m).
 - You can do this by Pythagoras: $(6370.01^2 - 6370^2)^{1/2} = 11.3$ km
 - However if we write $(6370 + 0.01)^2 \approx 6370^2 + 2 \times 0.01 \times 6370$, the 6370^2 cancels and we get $\sqrt{127.4} = 11.3$ km
 - this is much easier to check than the previous result (you can do $2 \times 63.7 = 127.4$ in your head, and since $11^2 = 121$, 11.3 must be about right)

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EXAMPLE 3: PRECISION

- The highest energy cosmic ray observed by the Auger experiment had a measured energy of 6.4 J. If the particle was a proton, mass 1.7×10^{-27} kg, calculate the difference between its velocity and the speed of light.
 - Relativity gives $E = \gamma mc^2$ where $\gamma = (1 - v^2/c^2)^{-1/2}$
 - so try this: $\gamma = 6.4 / (1.7 \times 10^{-27} \times (3.0 \times 10^8)^2) = 4.3 \times 10^{10}$
 - $1 - v^2/c^2 = 1/\gamma^2 = 5.5 \times 10^{-22}$
 - $v/c = (1 - 5.5 \times 10^{-22})^{1/2} = 1$ (exactly, on calculator), so $v \equiv c$
 - oops, can't be right, proton is not massless
 - Instead write $v = c(1 - \epsilon)$ where $\epsilon \ll 1$
 - $v^2 = c^2(1 - 2\epsilon)$ using binomial expansion
 - $\gamma = (2\epsilon)^{-1/2}$, so $\epsilon = 0.5/\gamma^2 = 2.7 \times 10^{-22}$
 - $c - v = c\epsilon = 8.2 \times 10^{-14} \text{ m s}^{-1}$ (!!)

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CHECKS AND ESTIMATES

- Easy to get numerical values wrong
 - mismatch units, e.g. 1.4 mm/184 m gives 7.6×10^{-3} rad instead of 7.6×10^{-6} rad if you forget to convert mm to m
 - think calculator's set to degrees when it's set to radians, or vice versa
 - forget square or square root (or indeed any other power)
 - transpose digits in calculator, e.g. 19 becomes 91
 - forget exponent, or reverse its sign, e.g. 1.6×10^{-19} becomes 1.6×10^{19}
- Therefore should always check before moving on
 - numerical mistakes are usually easy to fix *once you realise you have a problem*

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BASIC CHECKS

- Is the number obviously wrong?
 - satellite orbiting at $6 \times$ speed of light, stars in binary system separated by 32 cm, mass of snooker ball = 15 tonnes, etc.
 - Check for forgotten powers
 - Check for incorrect units
- Does the rest of the problem make sense with this number?
 - if you found that the mass of an object was 300 kg, and the question continues, "A student picks up the object" – your number is probably wrong!
- Is the number intrinsically plausible?
 - it is useful to be aware of "typical" values for particular areas of physics or astrophysics
 - e.g. stars typically between 0.1 and $100 \times$ Sun's mass

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ESTIMATES

- If at all doubtful, make estimate using powers of 10
 - Example: The Sun has a power output of 3.8×10^{26} W and a radius of 7.0×10^8 m. What is its surface temperature?
 - We use $L = 4\pi R^2 \sigma T^4$ where $\sigma = 5.67 \times 10^{-8}$ W m⁻² K⁻⁴
 - So we have $T = (L/4\pi R^2 \sigma)^{1/4}$
 - Estimate: $4 \times 10^{26} / (4 \times 3 \times (7 \times 10^8)^2 \times 6 \times 10^{-8})$
 - powers of 10: $26 - 2 \times 8 + 8 = 18$
 - numbers: $4 / (4 \times 3 \times 50 \times 6) = 1/900 = 10^{-3}$
 - so $T^4 \sim 10^{15}$ K⁴
 - $T = (10^{16}/10)^{1/4} = 10000 / (10^{1/4}) \sim 5000$ K (actually 5700 K)
 - Why do this instead of just redoing calculation?
 - it's quicker (with practice)
 - you may be entering calculation into calculator wrongly
 - if you focus on powers, you may find one you forgot!

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CALCULATIONS: SUMMARY

- Before entering numbers:
 - check dimensions of equation
 - check behaviour of equation (special cases etc.)
 - think about whether you should be using any expansions or approximations
 - check units of numerical values
- After entering numbers:
 - consider whether result is physically possible
 - consider whether it is plausible
 - if you made any approximations, check whether they were valid
 - e.g. if you assumed $v \ll c$, but the result is $v = 2.8 \times 10^8$ m s⁻¹, you have a problem

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PROBLEM SOLVING: SUMMARY

- Basic approach
 - **Model** the system, and identify the underlying physics
 - **Formulate** equations describing the system
 - **Solve** the equations for the variable(s) you need
 - **Check** that your solution makes sense before using it
- Principles
 - Use symbols, not numbers, for as long as possible
 - Always check dimensions
 - Look for special cases, and check solution works there
 - Check units before substituting numbers
 - Check that numerical answers are
 - physically possible
 - plausible for given system

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AND FINALLY...

PRACTISE!

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