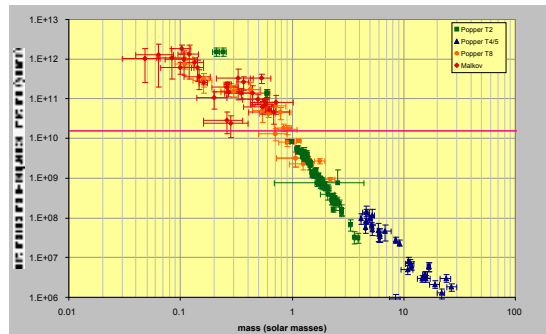


Star Birth

- Most of the bright stars we see have lifetimes much less than the age of the Solar System
 - star formation is an ongoing process
 - how does it happen?
 - are stars born now different from older stars?
 - what about planetary systems?

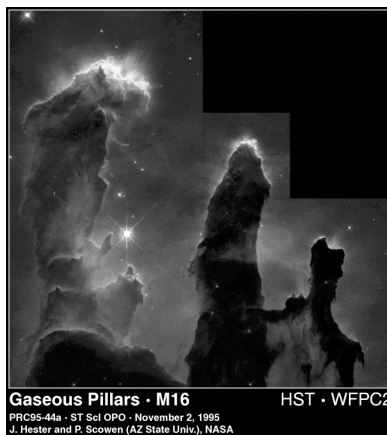


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How are stars born?



Gaseous Pillars • M16 HST • WFPC2
PRC95-44a • ST ScI OPO • November 2, 1995
 J. Hester and P. Scowen (AZ State Univ.), NASA

- Space is not empty, but filled with very rarefied gas
- Gas pressure depends on temperature: cool, dense gas may not have high enough pressure to balance inward gravitational force
- Cool, dense molecular gas can collapse to form stars

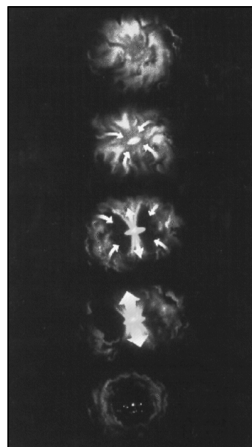
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Stages in star birth

- As gas cloud collapses
 - its rotation causes the formation of a disc around the young star
 - the gas forming the young star heats up as the star contracts
 - the increased pressure causes jets of gas to be emitted from the poles of the young star
 - conversion of gravitational energy to radiation (electromagnetic energy) causes young star to shine, even though fusion has not started
 - eventually fusion reactions turn on in the centre of the young star: it has now reached the main sequence



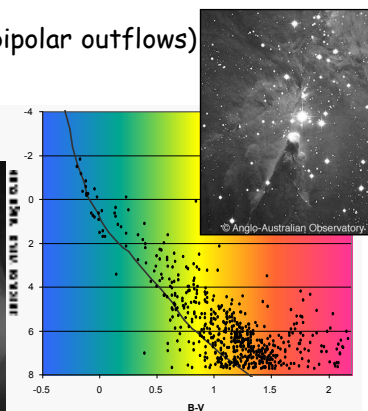
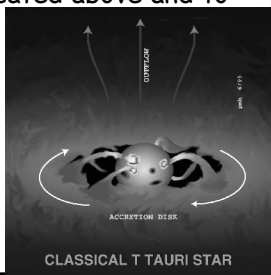
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Recognising young stars

- They will be surrounded by warm gas and dust, which gives off infra-red radiation
- They may emit jets of material (bipolar outflows)
- They are unstable, and hence may have variable light output
- They are located above and to the right of the main sequence (they are large and cool)

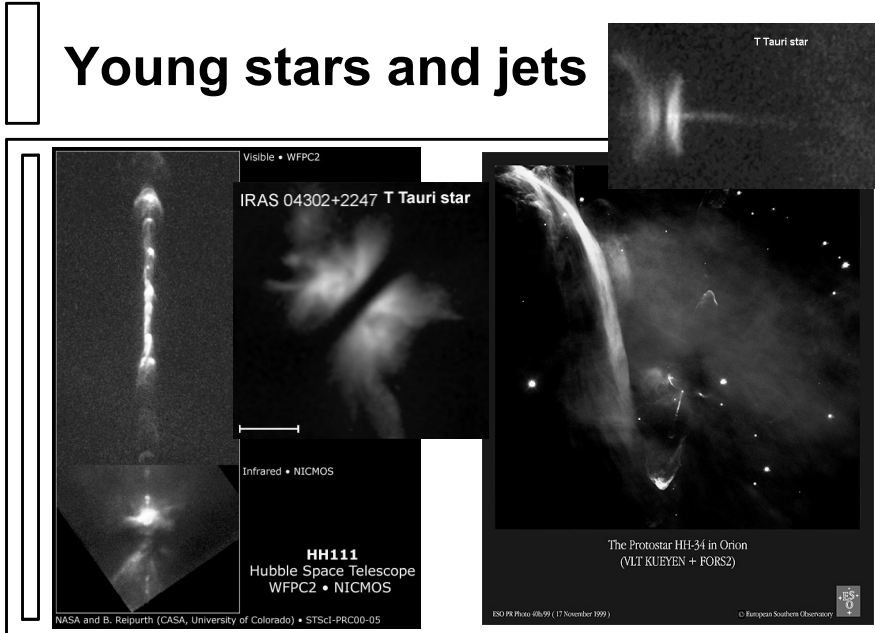


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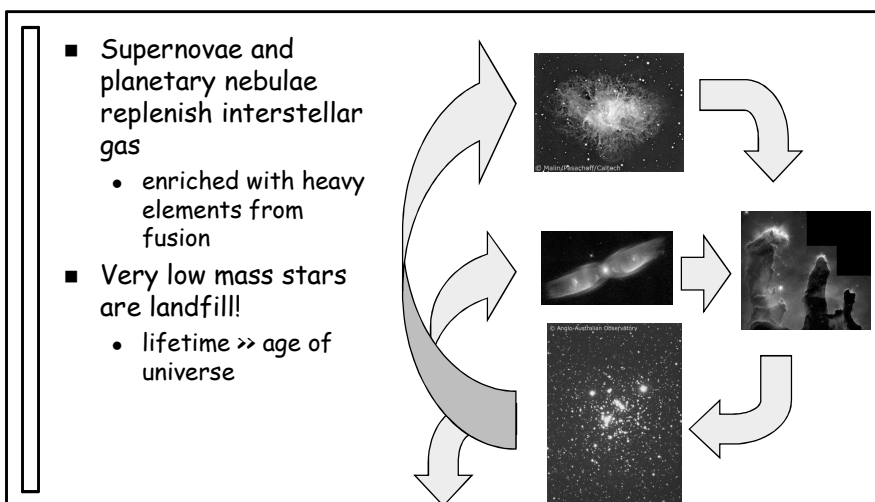
Young stars and jets



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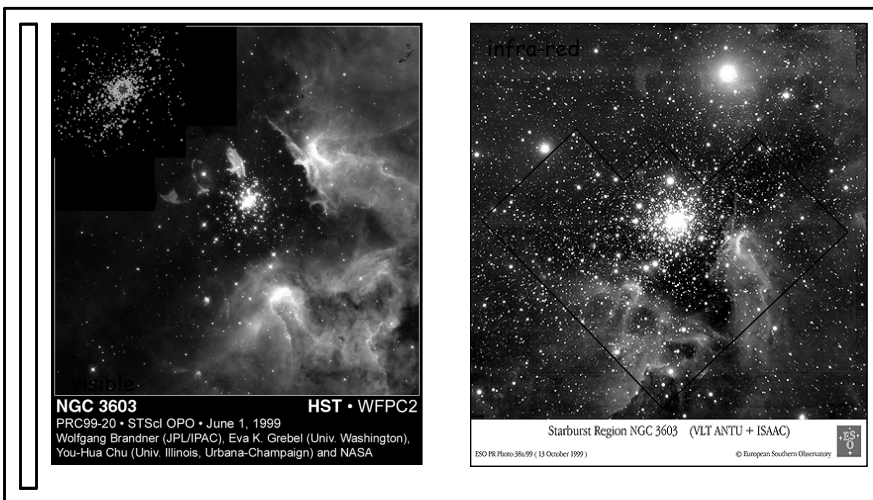
Galactic recycling



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A star formation region

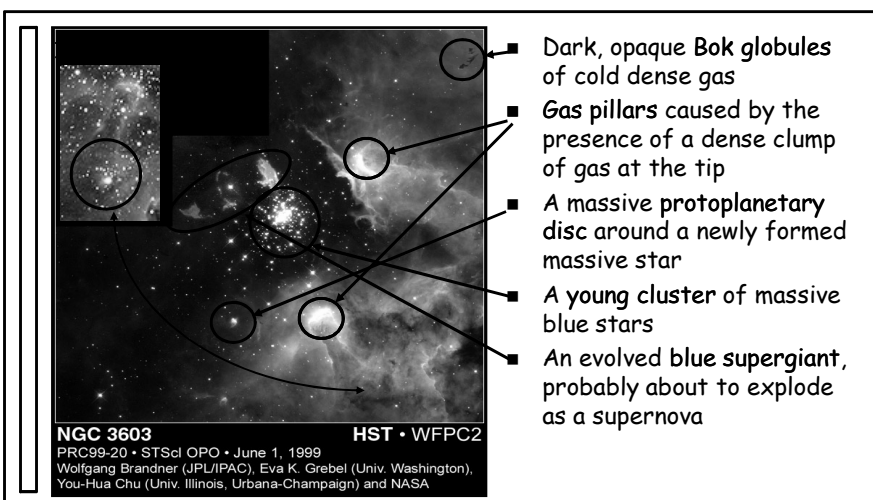


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Star formation in action: NGC3603



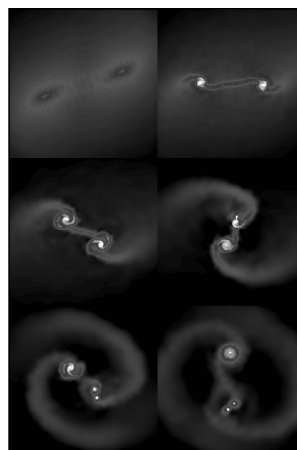
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Binary systems

- Most stars are members of binary or multiple systems
 - of 30 nearest stars, 12 are in binary systems and 6 in triples
- How does this happen?
 - clouds spin faster as they collapse
 - ◆ above critical speed cloud will break up into smaller clumps
 - ◆ these clumps form individual stars, still gravitationally bound together (since clump was)
 - picture shows simulation which produced triple system



Matthew Bate, U. of Exeter

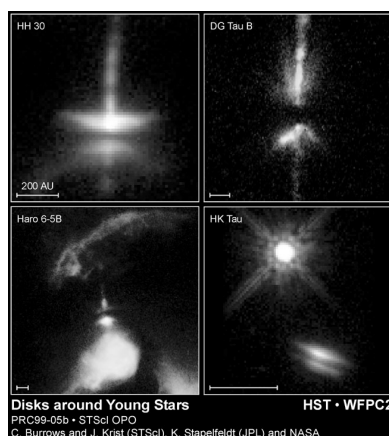
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Planetary systems

- Many young stars are seen to be surrounded by dusty discs ("protoplanetary discs" or "proplyds")
 - the Solar System is believed to have developed from such a disc
 - dust grains clump together to form larger objects, eventually planets
- Planetary systems common?



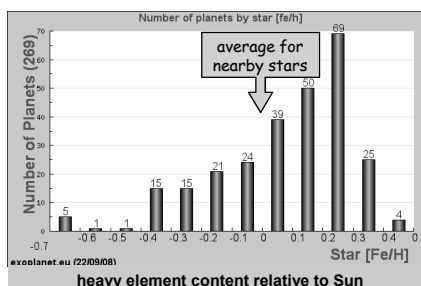
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Recycling and planets

- Many extrasolar planets have now been discovered (see later)
- Heavy element content of stars with planets systematically higher than typical sample, though some low-heavy-element stars do have planets
 - not surprising if planets form from build-up of dust grains into rocky bodies
 - implies planets commoner around second-generation stars formed from enriched material
 - indeed, no planets seen in transit search in globular cluster 47 Tucanae



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Brown dwarfs

- Stars must have masses at least 8% of the Sun's
 - otherwise fusion never starts (not hot enough)
- Jupiter is only 0.1% of Sun's mass
- Between the two are sub-stellar brown dwarfs
 - in theory brown dwarfs form "like stars", not "like planets"
 - observationally hard to define boundary!



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What have we learned?

- Star formation is an ongoing process
 - many of the bright stars we see are much younger than the Sun
- Stars form when a clump of dense, cool gas collapses under gravity
 - rapidly rotating clouds may fragment to produce binary systems
 - heavy-element-rich clouds yield stars with planets
- Supernovae and planetary nebulae recycle material to the interstellar gas
 - this is enriched in heavy elements
 - shocks from supernovae may also encourage stars to form
- Where is all this happening?
 - only in certain regions of our Galaxy

...next lecture!